## Newton-Raphson

- When the derivative of $f(x)$ is known, and when $f(x)$ is well behaved, the celebrated (and ancient) NewtonRaphson method gives the fastest convergence of all ("quadratic", i.e. $\mathrm{m}=2$, such that $\varepsilon_{\mathrm{n}}=\varepsilon_{\mathrm{n}-1}{ }^{2}$ )
- Relies on the Taylor expansion

$$
f(x+\delta)=f(x)+\delta f^{\prime}(x)+1 / 2 \delta^{2} f^{\prime \prime}(x)+\ldots \ldots
$$

If ith iteration, $x_{i}$, is close to the root, then for the next iteration, try $\mathrm{x}_{\mathrm{i}+1}=\mathrm{x}_{\mathrm{i}}+\delta$ with $\delta=-\mathrm{f}\left(\mathrm{x}_{\mathrm{i}}\right) / \mathrm{f}^{\prime}\left(\mathrm{x}_{\mathrm{i}}\right)$

## Geonnetric renresentation

- Newton-Raphson method: picture from Recipes
- Convergence is quadratic, with $\mathrm{m}=2$


Extrapolate the local derivative to find the next estimate of the root

## Newton-Raphson

- Convergence is rapid, and the method is very useful for "polishing" a root (i.e. refining an estimate that is nearly correct)

Clearly, a few iterations usually yields an accurate result in the limit of small $\delta$, because terms of order $1 / 2 \delta^{2} f^{\prime \prime}(\mathrm{x})$ or higher are much smaller than $\delta \mathrm{f}(\mathrm{x})$
$f(x+\delta)=f(x)+\delta f^{\prime}(x)+1 / 2 \delta^{2} f^{\prime \prime}(x)+\ldots .$.
Exception: when $\mathrm{f}^{\prime}(\mathrm{x})$ is very small (or zero)

## Newton-Raphson

- Cases where $f^{\prime}(x)$ is small (i.e. where there is a local extremum) can send the solution shooting off into outer space
.... unless the method is modified to keep the solution within a known bracket
(as in routine "RTSAFE")



## Summary of 1-D root finding

- Preferred methods:
- If $f^{\prime}$ is known analytically, "safe Newton-Raphson" is the preferred method
- If $\mathrm{f}^{\prime}$ is not known analytically, Brent's method is the method of choice
- Could use "safe Newton-Raphson" with the derivative computed numerically, but the latter takes up as much time as is saved by the use of $\mathrm{N}-\mathrm{R}$
- Newton-Raphson is also useful for rapid "polishing" (very close to the solution) and for multidimensional root finding.


## Roots of polynomial functions

- General polynomial equation

$$
0=P(x) \equiv \sum \mathrm{a}_{\mathrm{k}} \mathrm{x}^{\mathrm{k}}
$$

has an analytic solution for cases up to and including the quartic

No general analytic solution for quintics and higher ("the equation that couldn't be solved")

## Roots of polynomial functions

- General features:

Polynomial of order N will have N roots, which can be real or complex and may or may not be distinct

If the $a_{k}$ are all real, the roots may be real or complex, but any complex roots will occur in pairs of complex conjugates, $\mathrm{a} \pm \mathrm{bi}$

Polynomials can be ill-conditioned: a small change in the coefficients can cause a large change in the roots

## Roots of polynomial functions

- General features:

Degenerate roots, or nearly equal roots, present the biggest numerical problems

Example: $P(x) \equiv(x-c)^{2}=0$

- Brackets don't exist since $P(x)$ is never negative
- Derivative $f^{\prime}(x)=0$ at the solution $x=c$
$\rightarrow$ Newton-Raphson is potentially unstable


## Deflation

- Every time a root, $x_{k}$, is found, the order of the polynomial can be decreased by one, and the equation becomes
$0=P(x)=\left(x-x_{k}\right) Q(x)$
where $Q(x)$ can be obtained by "synthetic division" (see §5.3 in Recipes)
- This procedure, known as "deflation", allows the roots to be obtained one-by-one


## Laguerre's method for obtaining a single root

- Mathematical background

$$
\begin{aligned}
& P(x)=\left(x-x_{1}\right)\left(x-x_{2}\right) \ldots\left(x-x_{n}\right) \\
& \ln |P(x)|=\ln \left|x-x_{1}\right|+\ln \left|x-x_{2}\right|+\ldots .+\ln \left|x-x_{n}\right| \\
& G \equiv d \ln |P(x)| / d x=\left(x-x_{1}\right)^{-1}+\left(x-x_{2}\right)^{-1}+\ldots+\left(x-x_{n}\right)^{-1} \\
& H \equiv-d \ln |P(x)| / d x^{2}=\left(x-x_{1}\right)^{-2}+\left(x-x_{2}\right)^{-2}+\ldots+\left(x-x_{n}\right)^{-2}
\end{aligned}
$$

## Laguerre's method for obtaining a single root

- Now make a "drastic set of assumptions"

We suppose $x_{1}$ is located at distance a from our current guess, $x$, and that all the other roots lie at a distance $b$

$$
\begin{aligned}
& G=\left(x-x_{1}\right)^{-1}+\left(x-x_{2}\right)^{-1}+\ldots+\left(x-x_{n}\right)^{-1}=a^{-1}+(n-1) b^{-1} \\
& H=\left(x-x_{1}\right)^{-2}+\left(x-x_{2}\right)^{-2}+\ldots+\left(x-x_{n}\right)^{-2}=a^{-2}+(n-1) b^{-2}
\end{aligned}
$$

We can then eliminate $b$ from the above equations to obtain a quadratic equation for $\mathrm{a}^{-1} \ldots .$.

## Laguerre's method for obtaining a single root

## We then obtain

$$
a=\frac{n}{G \pm \sqrt{(n-1)\left(n H-G^{2}\right)}}
$$

This, of course, is not the exact value of $\left(x-x_{1}\right)$, because of the approximation made in obtaining it, but it is useful for obtaining the next estimate of $x_{1}$, which we take as ( $\mathrm{x}-\mathrm{a}$ )

## Overview

- Find roots one at a time, using Laguerre's method iteratively until $a$ is sufficiently small
- After each root is found, reduce the order of the polynomial by deflation.


## Multidimensional non-linear systems of equations

- In general, a horrible problem, for reasons shown schematically in Recipes, Fig 9.6.1



## Multidimensional non-linear systems of equations

- The only general algorithm is Newton-Raphson

Generalization to multidimensional case $\mathrm{f}(\mathrm{x}+\delta)=\mathrm{f}(\mathrm{x})+\delta \mathrm{f}^{\prime}(\mathrm{x})+(1 / 2) \delta^{2} \mathrm{f}^{\prime \prime}(\mathrm{x})+\ldots .$.
becomes
$\mathbf{f}(\mathbf{x}+\boldsymbol{\delta})=\mathbf{f}(\mathbf{x})+\mathbf{J}(\mathbf{x}) \delta+O\left(\delta^{2}\right)$
where $\mathbf{J}$ is the Jacobian matrix, with $J_{i k}=\partial f_{i} / \partial x_{k}$

## Multidimensional non-linear systems of equations

The next iteration on $x$ is therefore

$$
\mathbf{x}^{\prime}=\mathrm{x}+\delta=\mathrm{x}-\mathrm{J}^{-1} \mathrm{f}(\mathrm{x})
$$

where $\delta=-\mathbf{J}^{-1} \mathbf{f}(\mathbf{x})$ is obtained by the solution of the linear set of equations $\mathbf{J} \delta=-\mathbf{f}$
(which we know how to do! ©)

As in 1-D, this only works with a smooth function or a good first guess

## Newton-Raphson with a modified step size

- Can improve the robustness of multidimensional $\mathrm{N}-\mathrm{R}$ by considering the quantity:
$Q \equiv 1 / 2|f|^{2}=1 / 2 \sum f_{i} f_{i}$
which tends to zero at the root

Any good step should make Q smaller

## Newton-Raphson with a modified step size

- Consider now the gradient of Q

$$
\begin{aligned}
& \partial \mathrm{Q} / \partial \mathrm{x}_{\mathrm{j}}=\partial\left(1 / 2|\mathrm{f}|^{2}\right) / \partial \mathrm{x}_{\mathrm{j}}=\sum \mathrm{f}_{\mathrm{i}} \partial \mathrm{f}_{\mathrm{i}} / \partial \mathrm{x}_{\mathrm{j}} \\
\Rightarrow & \nabla \mathrm{Q}=\mathbf{f}^{\top} \mathbf{j}
\end{aligned}
$$

Our step $\delta=\mathbf{J}^{-1} \mathbf{f}$ is in the right direction to reduce $Q$, because

$$
\nabla Q . \delta=-\mathbf{f}^{\top}\left(J^{-1}\right) f=-2 Q<0
$$

## Newton-Raphson with a modified step size

So if adding $\delta$ to $\mathbf{x}$ doesn't decrease $Q$, a sufficiently small step in the same direction must

Strategy: if the solution starts to overshoot - as indicated by monitoring of $Q \equiv 1 / 2|f|^{2}$ - use a smaller step $\lambda \delta$ in the same direction (where $\lambda$ is a positive scalar <1)

See Recipes $\S 9.7$ for the details of how to choose $\lambda$

This is implemented in the function NEWT, which will successfully find a root for almost any reasonable initial guess.

## Example application: interstellar chemistry

- More than one hundred different molecules have been detected in the interstellar gas
- They are formed and destroyed by a complex network of reactions involving bimolecular reactions and unimolecular processes such as photodissociation


## Example application: interstellar chemistry

- Consider the reaction $A+B \rightarrow C+D$ :

This destroys $A$ and $B$ (and creates $C$ and $D$ ) at a rate $k_{r} n(A) n(B)$ per unit volume, where $n(X)$ is the density of species $X$ (number per unit volume) and $k_{r}$ is the rate coefficient (units: $\mathrm{cm}^{3} \mathrm{~s}^{-1}$ ) for the reaction in question

- Molecule A might also be destroyed by photodissociation, $A+h \nu \rightarrow E+F$
This destroys $A$ (and creates $E$ and $F$ ) at a rate $\zeta_{p} n(A)$ per unit volume, where $n(X)$ is the density of species $X$ (number per unit volume) and $\zeta_{p}$ is the photodissociation rate (units: $\mathrm{s}^{-1}$ ) for this process


## Rate equations for interstellar chemistry

- The density of the $i$ th molecule therefore obeys the rate equation:

Processes that create ith molecule Reactions that destroy ith molecule $\partial n_{i} / \partial t=\sum_{j} \sum_{k} k_{i j k} n_{j} n_{k}+\sum_{k} \zeta_{i k} n_{k}-n_{i} \sum_{m} \sum_{k} k_{m i k} n_{k}-n_{i} \sum_{m} \zeta_{m i}$

Total rate at which $k$ undergoes unimolecular processes to produce i

## Equilibrium solution

- In equilibrium, the molecular densities will obey $\mathbf{f}(\mathbf{n})=\mathbf{0}$
where $\mathrm{n}_{\mathrm{i}}$ is the density of molecule number $i$, and $\mathrm{f}_{\mathrm{i}} \equiv \partial \mathrm{n}_{\mathrm{i}} / \partial \mathrm{t}$ is a quadratic function of the $\mathrm{n}_{\mathrm{i}}$

A multidimensional root-finding problem!
... and one for which the Jacobian is easily computed (quadratic function)

$$
\begin{array}{ll}
\partial \mathrm{f}_{\mathrm{i}} / \partial \mathrm{n}_{\mathrm{k}}=\sum_{\mathrm{j}} \mathrm{k}_{\mathrm{ijk}} \mathrm{n}_{\mathrm{j}}+\zeta_{\mathrm{ik}}-\mathrm{n}_{\mathrm{i}} \sum_{\mathrm{m}} \mathrm{k}_{\mathrm{mik}} & (\mathrm{i} \neq \mathrm{k}) \\
\partial \mathrm{f}_{\mathrm{i}} / \partial \mathrm{n}_{\mathrm{i}}=-\sum_{\mathrm{m}} \mathrm{k}_{\mathrm{mik}} \mathrm{n}_{\mathrm{k}}-\sum_{\mathrm{m}} \zeta_{\mathrm{mi}} & \text { (diagonal elements) }
\end{array}
$$

Newton's method generally works well

## Complication

- The equilbrium equations, as written, are singular, since they don't uniquely specify the total density of molecules We need to apply some constraints of the form

$$
\sum \mathrm{n}_{\mathrm{i}} \mathrm{~N}_{\mathrm{ic}}=\mathrm{n}\left(\mathrm{E}_{\mathrm{c}}\right)
$$

$\mathrm{N}_{\mathrm{ic}}=$ number of atoms of element C contained in molecule i
$\mathrm{n}\left(\mathrm{E}_{\mathrm{c}}\right)=($ fixed $)$ density of C nuclei

## Non-equilibrium solution

- In astrochemistry, the timescale for reaching equilibrium can sometimes be long compared to the ages of molecular clouds
Therefore, it is also interesting to integrate the equations

$$
\partial \mathrm{n}_{\mathrm{i}} / \partial \mathrm{t}=\sum_{\mathrm{j}} \sum_{\mathrm{k}} \mathrm{k}_{\mathrm{ijk}} \mathrm{n}_{\mathrm{j}} \mathrm{n}_{\mathrm{k}}+\sum_{\mathrm{k}} \zeta_{\mathrm{ik}} \mathrm{n}_{\mathrm{k}}-\mathrm{n}_{\mathrm{i}} \sum_{\mathrm{m}} \sum_{\mathrm{k}} \mathrm{k}_{\mathrm{mik}} \mathrm{n}_{\mathrm{k}}-\mathrm{n}_{\mathrm{i}} \sum_{\mathrm{m}} \zeta_{\mathrm{mi}}
$$

(a coupled set of ODE's, to be considered later)

In the limit of large $t$, this tends to the equilibrium solution
$\rightarrow$ an alternative method even if we are uninterested in the time evolution

